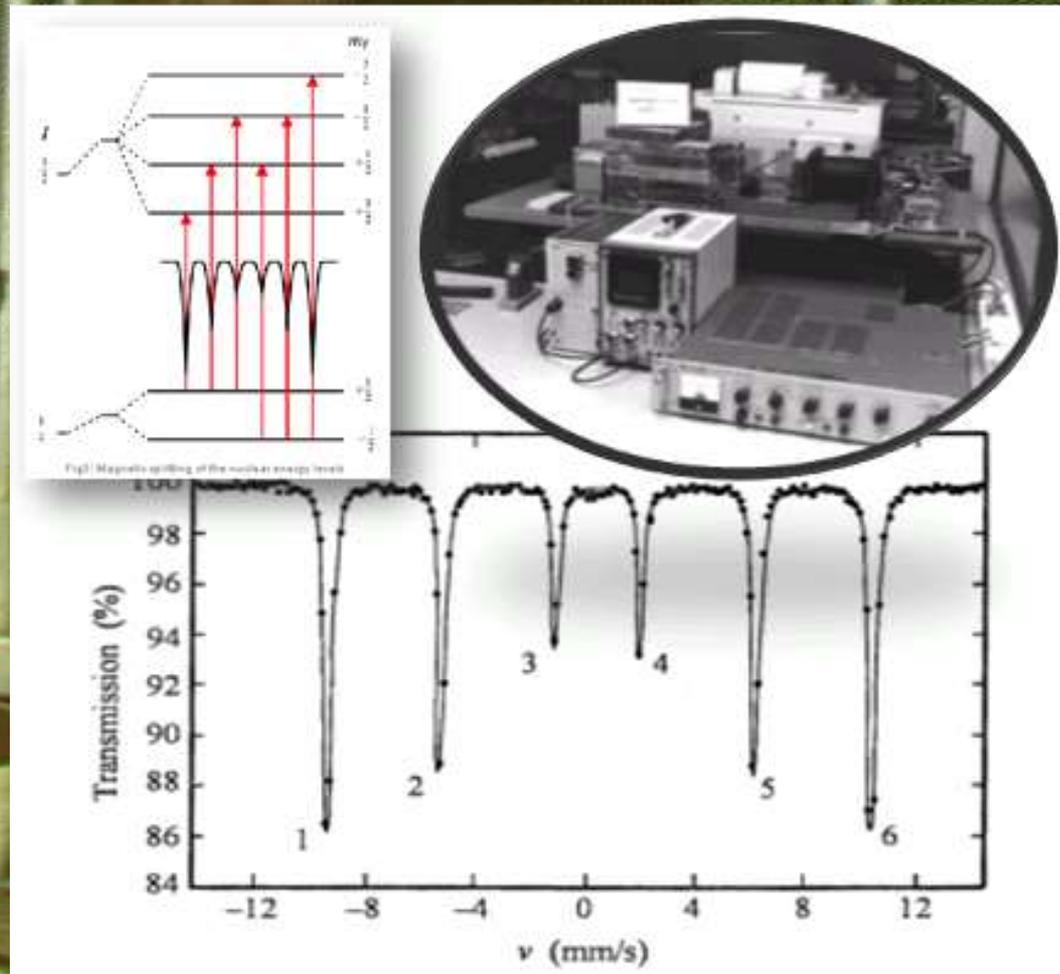


Mössbauer Spectroscopy II



Agenda: ANSEL Mössbauer Experiment

Mössbauer (Mössbauer) Spectroscopy with proportional counters:

Ultra-high-precision photon energy measurement:
Precision scanning resonant-absorption spectroscopy
with doppler-shifted photon energy, using gas amplification counters.

➤ Gas amplification counters, proportional counters, electronics.

➤ Mössbauer Principles:

Resonant γ absorption.

Recoil effects in γ emission and absorption,

Recoilless γ absorption by macroscopic samples,

➤ Determination of electric and magnetic HF interactions in various chemical Fe compounds

Reading Assignments:

(Knoll, LN): X ray spectroscopy with proportional counters (PC),

E_γ -dependent absorption coefficients, gas amplification counters,

Response of proportional counters to γ - and X rays, spurious peaks.

The Mössbauer Effect



1961 Nobel Prize in Physics.

Discovered (1958) recoilless nuclear fluorescence of gamma rays in ^{191}Ir .

Famous application: proof of red shift of gamma radiation in the gravitational field of the earth (Robert Pound and Glen Rebka);

Pound-Rebka experiment was one of the first experimental precision tests of Albert Einstein's theory of general relativity.

Long-term importance:

Use of Mössbauer effect in "Mössbauer spectroscopy" testing solid-state and chemical environments via electric and magnetic hyperfine interactions between atomic electrons and nuclear charge and magnetization distributions.

Nuclear Resonance Photo Absorption

Absorption of radiation = competition of various interactions between photons and microscopic structure (atomic, nuclear) of material →

absorbance = sum of statistical probabilities per constituent.

→ Mass absorption coefficient (photons): $\mu(E_\gamma) = \sum_n \mu_n(E_\gamma)$ ←

^{57}Co

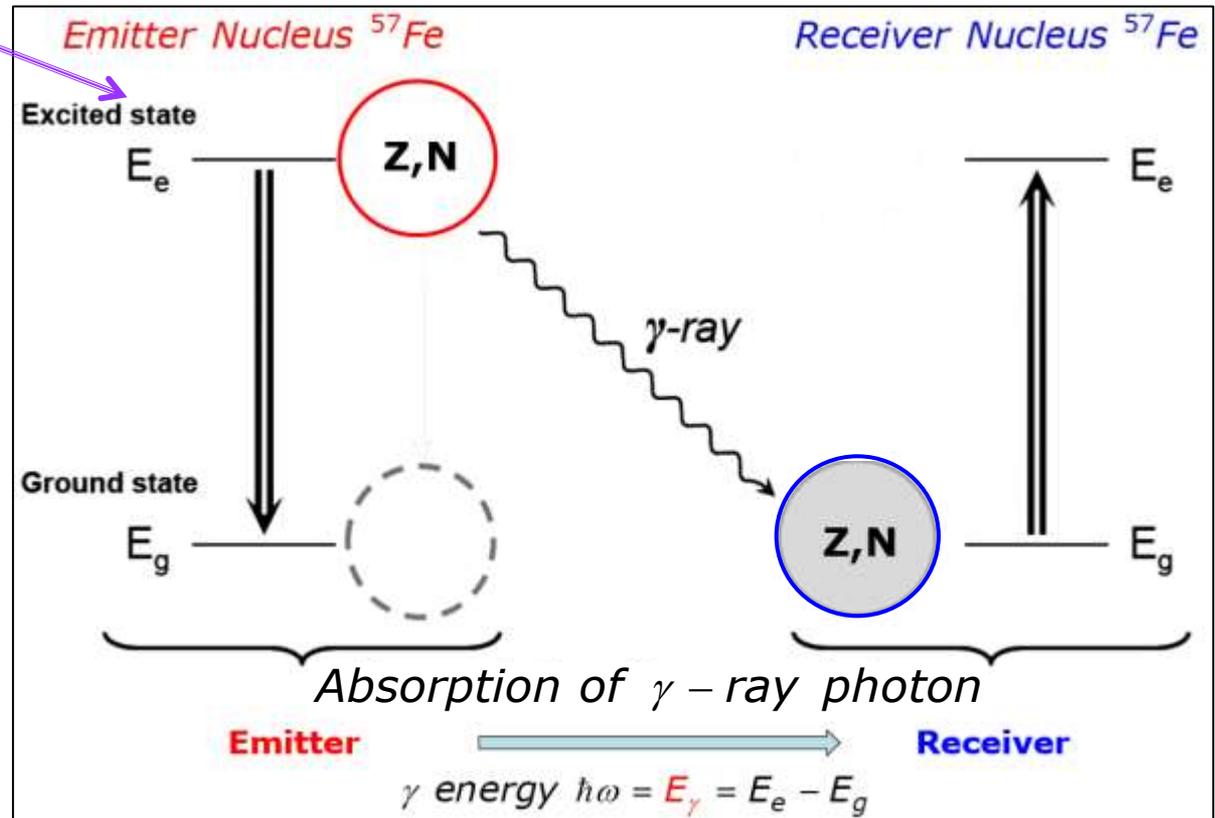
Nuclear Resonance Absorption
 $\mu_\nu(E_\gamma) = \mu_{\text{res}}(E_\gamma)$

Breit – Wigner Resonance

$$\mu_{\text{res}}(E_\gamma) \propto \frac{(\Gamma_e/2)^2}{(\Gamma_e/2)^2 + (E_\gamma - E_e)^2}$$

Resonance parameters E_e, Γ_e

→ Free nucleus:
Recoil effects
 $E_\gamma \neq E_e - E_g$



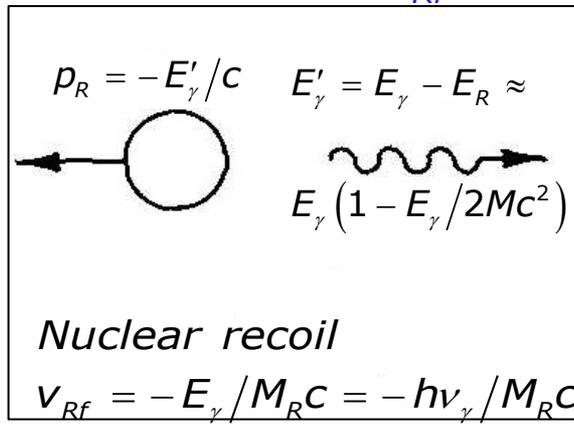
Spectroscopy Challenge: Doppler Shifted γ Energy

Emission and absorption of γ -rays by nuclei in motion (v thermal lattice vibrations, recoil effects due to γ emission) \rightarrow **Doppler effect** both in emission and absorption.

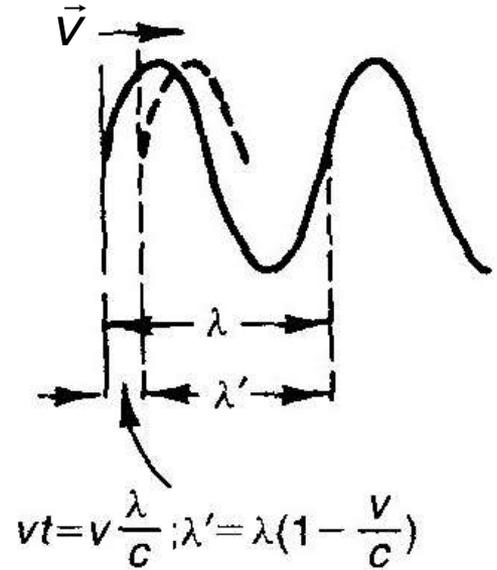
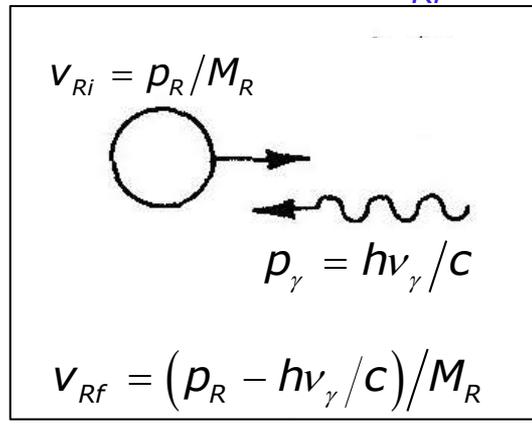
\rightarrow γ emission or absorption energy is different from

Nominal transition energy $h\nu_{if} = (E_i - E_f)$ (photon and nuclear recoil) $E_\gamma(v) \neq E_i - E_f$

γ emission @ $v_{Ri} = 0$

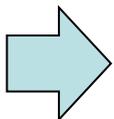


γ absorption @ $v_{Ri} \neq 0$



Typical Lab recoil velocities $v_R \sim (10^3 - 10^5) \text{ cm/s}$

Observer Lab velocity v
 $\rightarrow \lambda$ red(blue) shift



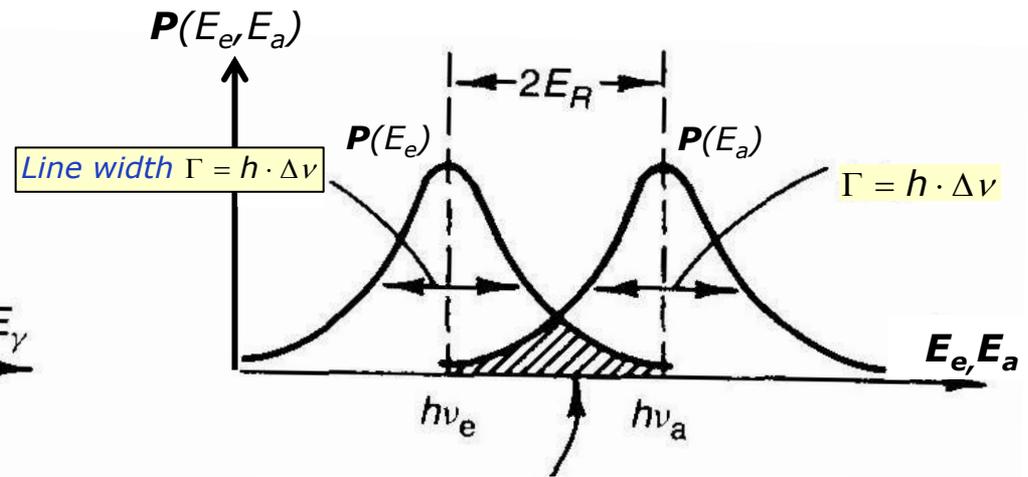
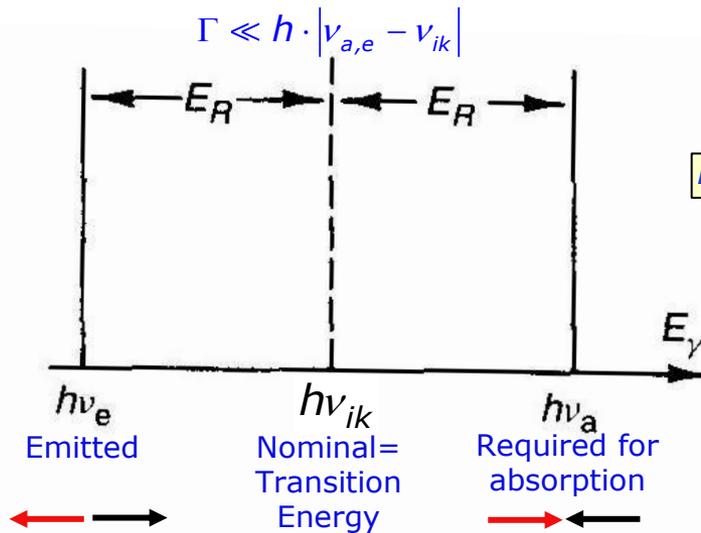
Doppler Shift : $E_\gamma(v) \approx hv_{ik} \cdot (1 \pm v/c)$

\pm sign : v blue or red shift

Emission/Absorption of γ -Rays in Thermal Environment

Quantum Effect: System absorbs electromagnetic radiation strongly if γ energy $h\nu$ equals a system energy level difference: $h\nu_{ik} = (E_i - E_k)$. For absorption, the lower level (i, k) must be occupied, the other empty.

→ Use for scanning level scheme $\{E_n\}$



Transition can occur only in overlap region

Quantum Effect → Coherent line broadening.

Normal environment $T \neq 0$ →

Thermal motion of nuclei → Incoherent line broadening.

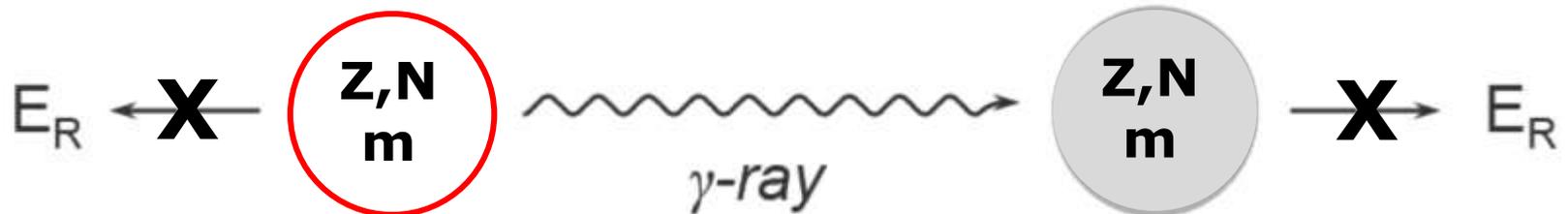
Velocity distributions of emitters and absorbers lead to broad line shapes, wash out resonance requirement, broader with increasing T .

Recoilless Emission/Absorption

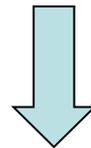
Momentum-energy transfer to nucleus (mass m) changes effective γ energy \rightarrow Loss of resonance condition

Emitter Nucleus ^{57}Fe

Receiver Nucleus ^{57}Fe

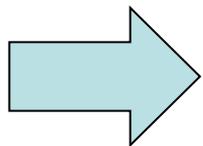


Recoil energy $E_R \approx \frac{E_\gamma^2}{2mc^2}$



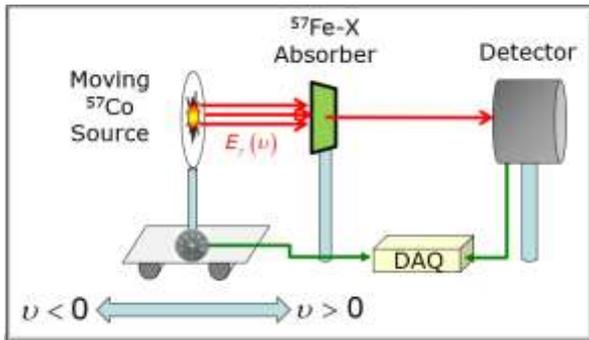
Anchored to macro crystal

Recoil energy $E_R = \frac{E_\gamma^2}{2M} \approx 0$



Momentum-energy transfer to nucleus embedded in macro crystal lattice is negligible \rightarrow Resonance condition retained \rightarrow Allows for precision absorption/emission spectroscopy!

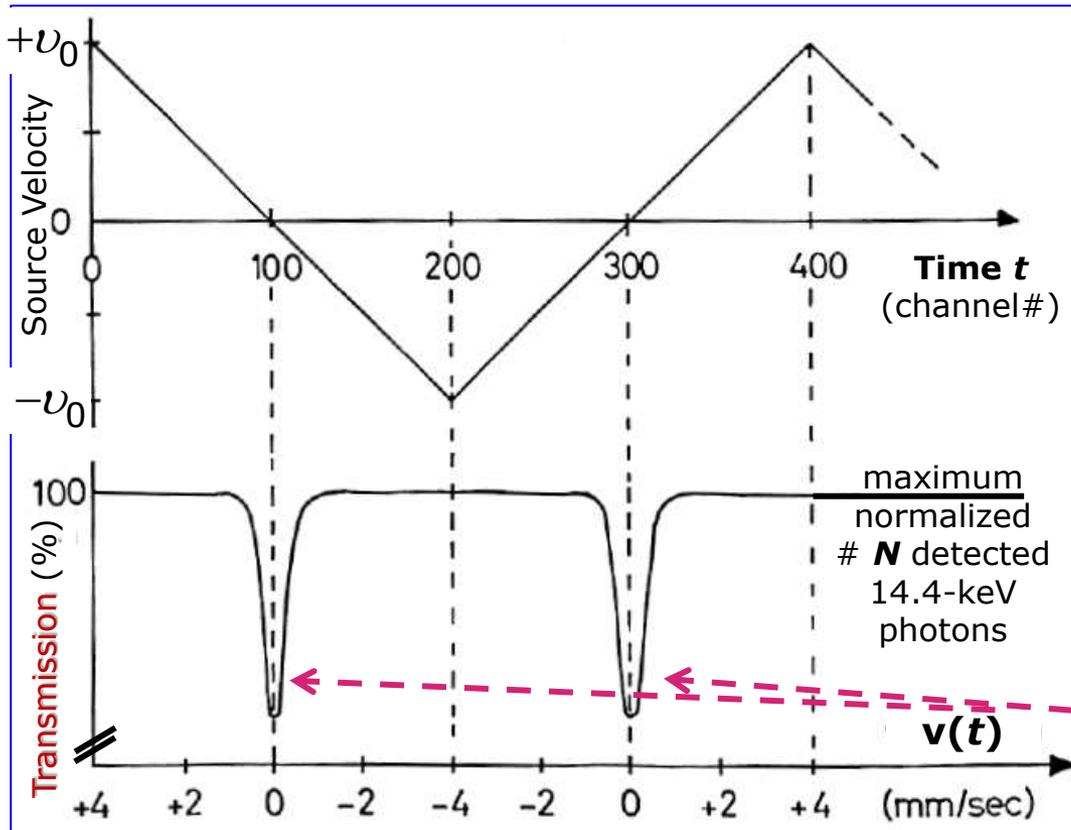
Precision Absorption Spectroscopy



$$\text{Tunable } E_\gamma(v) = 14.4(1 \pm v/c) \text{ keV}$$

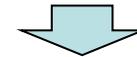
Lower transmission for $E_\gamma(v) = E(\langle {}^{57}\text{Fe} \rangle)$

Calibration $\langle {}^{57}\text{Fe} \rangle \equiv {}^{57}\text{Fe}$ in Fe absorber lattice



Example:

Identical crystal lattices of source & absorber and *linear velocity* drive. No external/crystal fields



$$v(t) = v_0 \cdot (t_0 \pm t); \text{ linear}$$

$$t \in [nT, (n+1)T]$$

$$\rightarrow dN/dt \propto dN/dv$$

Maximum (resonant) no-field absorption at $E(\langle {}^{57}\text{Fe} \rangle) = E_\gamma(v=0)$

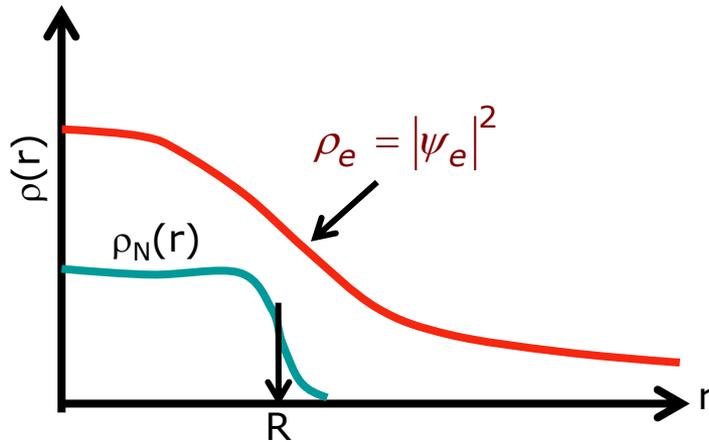
Mössbauer Spectroscopy Applications

Precise level energy scan, resolution $\Delta E \sim 10^{-10}$ eV

→ Investigate small perturbations of nuclear level energies due to interactions between

- ❑ Nuclear charge distributions and electronic density distributions in molecules, solid lattices
→ chemical shifts, electrostatic hyperfine interactions;
- ❑ Nuclear spins and magnetic moments with external magnetic fields, man-made or in lattices
→ spin and g-factor determinations, magnetic hyperfine interactions.

Isomer (Chemical) Shift of Atomic States



Perturbation theory calculation of nuclear energy level, **perturbation = H'** due to interaction of $\rho(r)$ with electrons $\psi(r)$

$$H'(r) = V(r < R) - V_0(r) = \text{interaction}$$

$$\delta E_e = \langle \psi_e | H' | \psi_e \rangle \equiv \int \psi_e^*(r) H'(r) \psi_e(r) d^3r$$

$$\delta E_e = \frac{1}{10\epsilon_0} Z e^2 R_n^2 |\psi_e(0)|^2 \quad \text{Nuclear state (n=0,1,...)}$$

$$\Delta E = \frac{1}{10\epsilon_0} Z e^2 |\psi_e(0)|^2 (R_1^2 - R_0^2) \quad \text{Difference between 2 states}$$

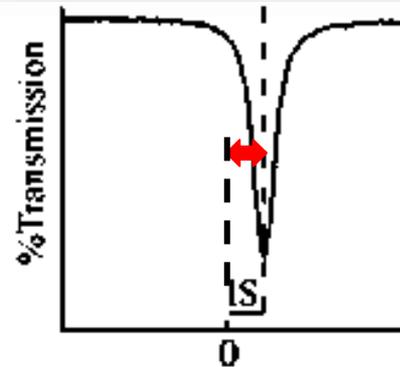
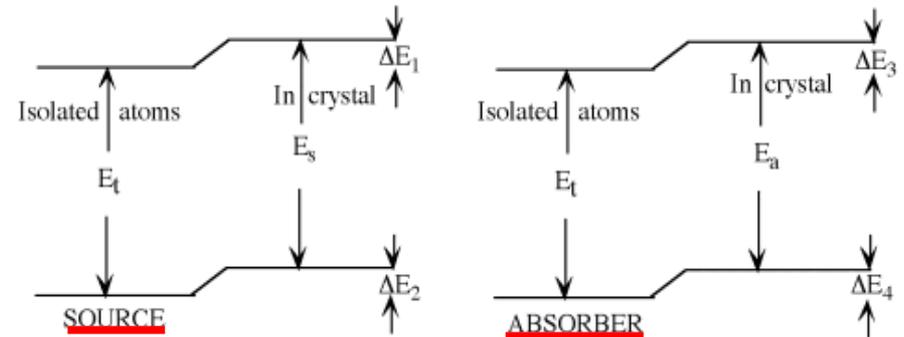
$$\begin{aligned} \Delta E_{\text{isomer}} &= \Delta E_{\text{absorb}} - \Delta E_{\text{source}} = \\ &= \frac{1}{10\epsilon_0} Z e^2 (R_1^2 - R_0^2) \left[|\psi_e^{\text{absorb}}(0)|^2 - |\psi_e^{\text{source}}(0)|^2 \right] \end{aligned}$$

Coulomb potential for spatially extended nucleus \rightarrow depends on R

Point \rightarrow finite size \rightarrow **Perturbation** $H'(r) = V(r) - V_0(r)$

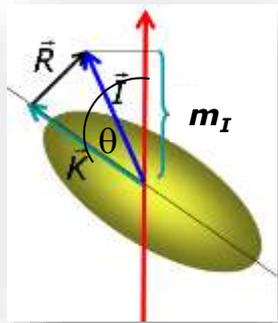
$$V_0 = \frac{1}{4\pi\epsilon_0} \frac{Ze}{r} \quad \text{for } r \geq R_n$$

$$V(r < R) = \frac{Ze}{4\pi\epsilon_0 R} \left(\frac{3}{2} - \frac{1}{2} \left(\frac{r}{R} \right)^2 \right) \quad \text{Finite Size nuclear states Radii } R=R_n \text{ (n=0,1,...)}$$



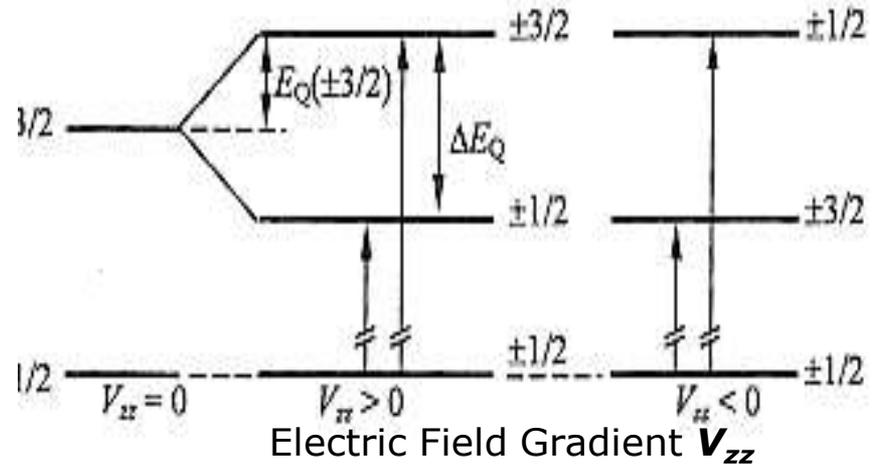
Transmission of γ -rays through absorber depends on source velocity \rightarrow scan with $T=T(v)$

Electric Quadrupole Hyperfine Interaction in Atoms

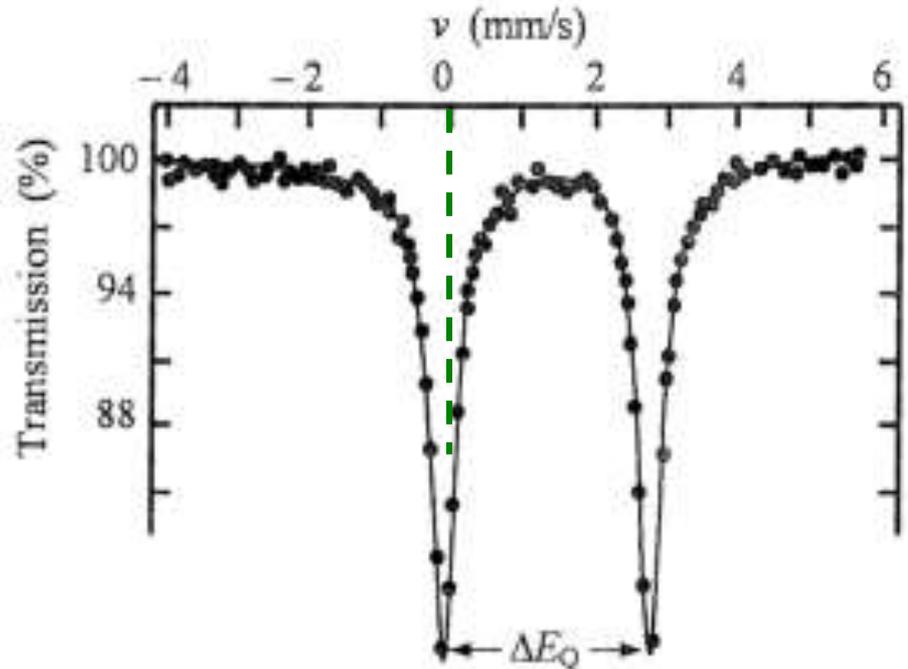


Nuclear electric quadrupole moment eQ measures deviation of nucleus from sphere. $Q_{\text{eff}}(I)$ can be aligned via interactions of external fields (spin I alignment).

Energy shift depends on orientation of Q (i.e., I) with respect to crystal field gradient. $Q_{\text{eff}} = Q' = 0$ for $I = 0, 1/2$



Electric Field Gradient V_{zz}



$$eQ = \int \rho(r) r^2 (3 \cos^2 \theta - 1) d^3r$$

$$E_Q = \frac{1}{4} eQ' \left(\frac{3}{2} \cos^2 \theta - \frac{1}{2} \right) \left(\frac{\partial^2 V_{\text{ext}}}{\partial z^2} \right)_{z=0}$$

Orientation (I , m_I) dependent quadrupole shift

$$E_Q = eQ \frac{3m_I^2 - I(I+1)}{4I(2I-1)} \left(\frac{\partial^2 V_{\text{ext}}}{\partial z^2} \right)_{z=0}$$

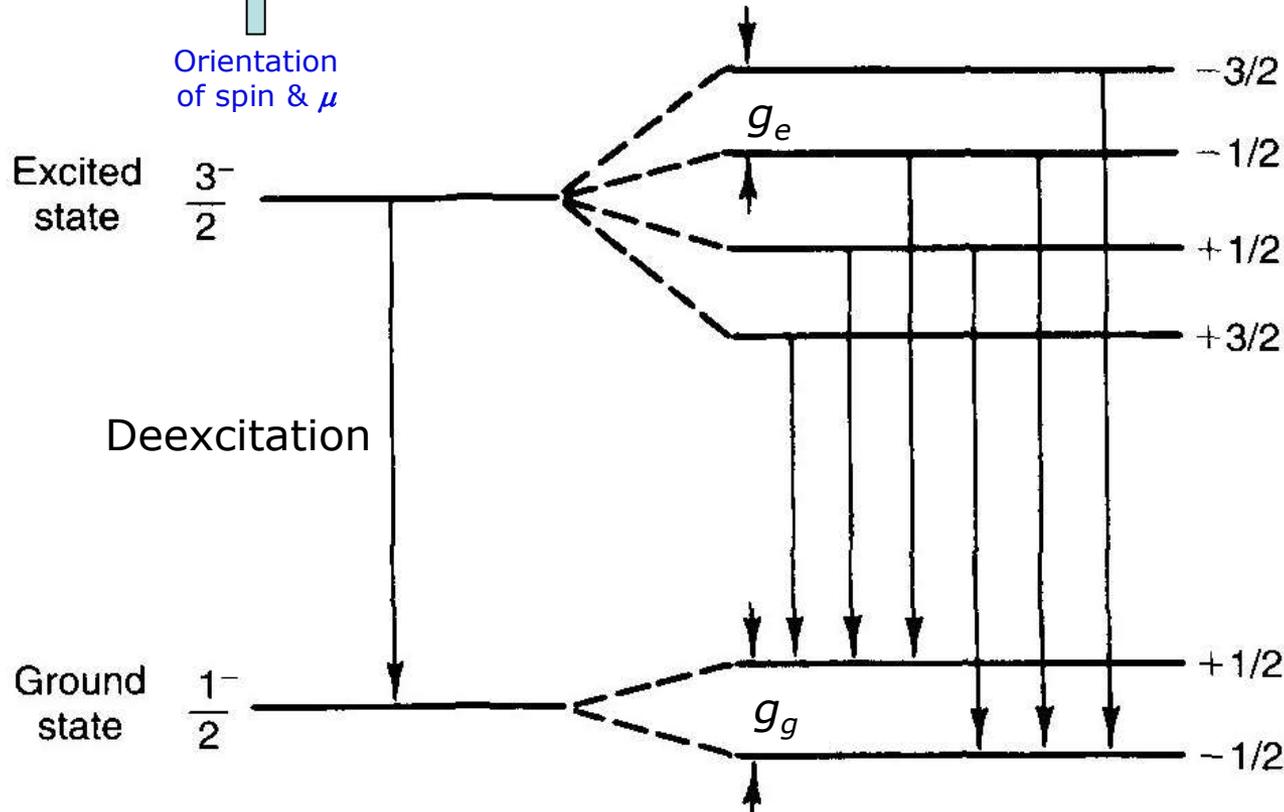
Magnetic Hyperfine Splitting in Atoms

Nuclear magneton $\mu_N = 5.05078324(13) \times 10^{-27} \text{ J/T}$

$$\Delta E_n = -\vec{\mu} \cdot \vec{B} = -m_I g_n \mu_N B_z$$

↑
Orientation
of spin & μ

Nuclear magnetic moment μ oriented || to spin I , can be aligned w/r to magnetic field B . g_N = gyro-magnetic ratio: different for different nuclear states → structure information!



Hyperfine structure splitting of the nuclear energy levels of ^{57}Fe . (a) When stainless steel is used, the levels are not split. (b) In ordinary iron, however, both levels are split, giving rise to a hyperfine structure with six components.

Ferro-Magnetic HF Interaction

$$\delta E_i^{\text{exp}} = E_\gamma \left(1 + \frac{v_1}{c} \right) \text{scan}$$

HF : $E_{\vec{\mu}} = -\vec{\mu} \cdot \vec{B} = -m_I \cdot (g \cdot \mu_N)_I \cdot |\vec{B}|$
 Nuclear states $\mu_{g,e} = +(I \cdot g)_{g,e} \mu_N$

$g_{e,g}$ = gyro-magnetic ratios
 μ_N = Bohr Magneton (unit)

$$\delta E_{\Delta\mu_{ij}} = -(g_e m_i - g_g m_j) \mu_N B$$

Chemical Shift IS

$$\begin{aligned} \delta E_1 &= E'_\gamma - (3/2q_e - 1/2q_g)\mu_N B \\ &= E'_\gamma - (\mu_e - \mu_g)B \end{aligned}$$

$$E_\gamma \left(1 + \frac{v_1}{c} \right) = E'_\gamma - (\mu_e - \mu_g)B$$

$$E_\gamma \left(1 + \frac{v_2}{c} \right) = E'_\gamma - (\mu_e/3 - \mu_g)B$$

$$E_\gamma \left(1 + \frac{v_3}{c} \right) = E'_\gamma - (-\mu_e/3 - \mu_g)B$$

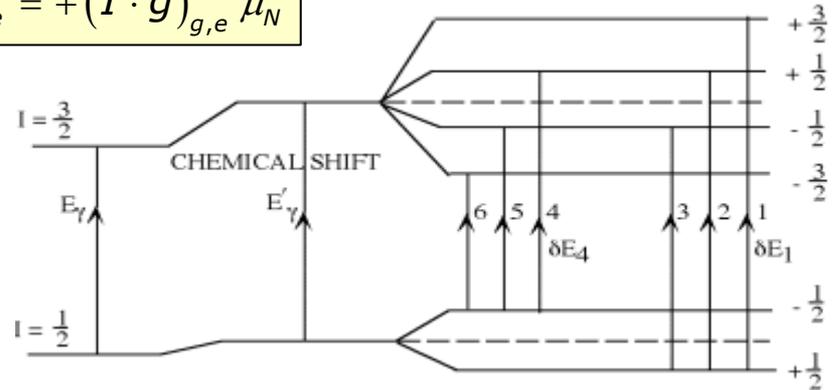
$$E_\gamma \left(1 + \frac{v_4}{c} \right) = E'_\gamma - (\mu_e/3 + \mu_g)B$$

$$E_\gamma \left(1 + \frac{v_5}{c} \right) = E'_\gamma - (-\mu_e + \mu_g)B$$

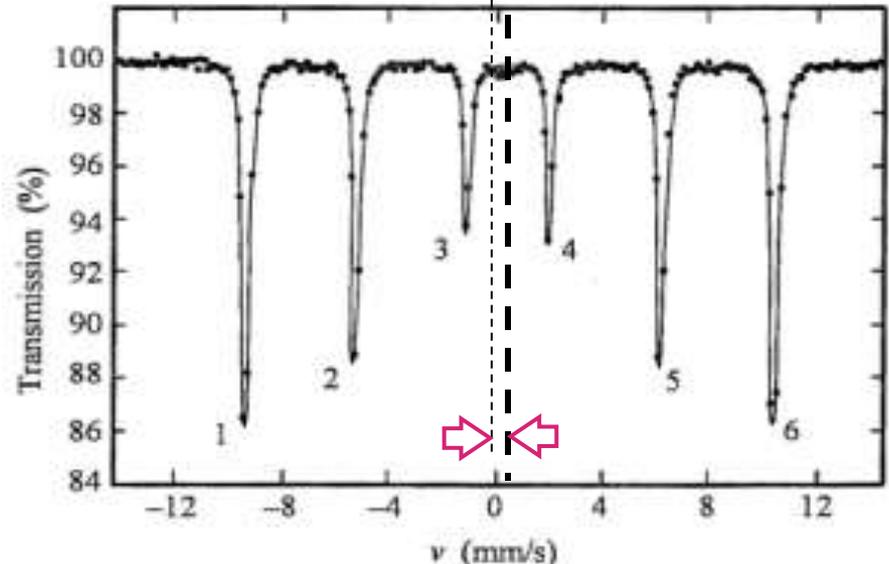
$$E_\gamma \left(1 + \frac{v_6}{c} \right) = E'_\gamma - (-\mu_e - \mu_g)B$$

$$(v_1 - v_2) \frac{E_\gamma}{c} = -\frac{2}{3} \mu_e B$$

$$(v_2 - v_4) \frac{E_\gamma}{c} = 2\mu_g B \quad \Rightarrow \quad \frac{v_1 - v_2}{v_2 - v_4} = \frac{-\mu_e}{3\mu_g}$$



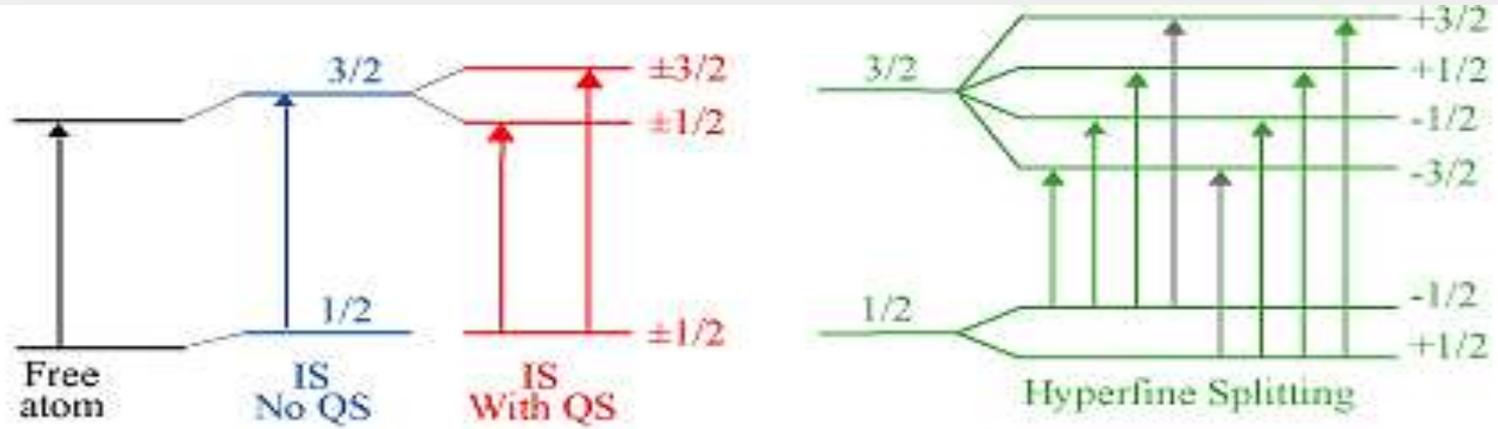
Ferromagnetic Fe absorber → observe magnetic HF



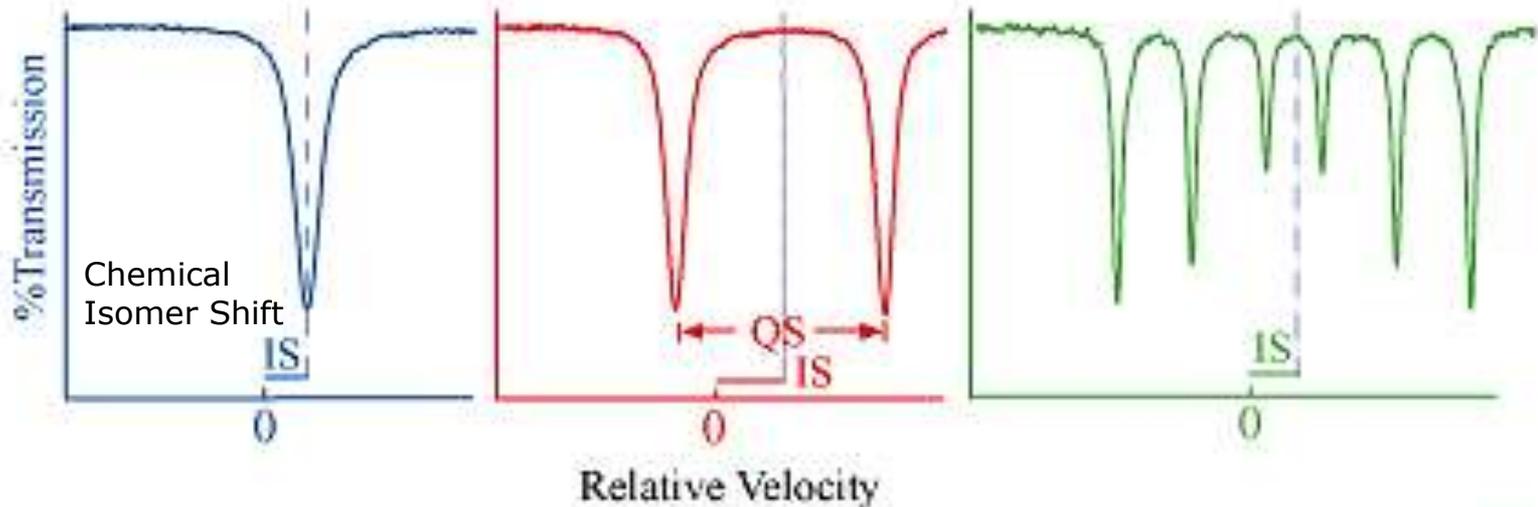
Velocity $v = -v_1 < 0$ source moving towards absorber

Electric + Magnetic HF Interactions

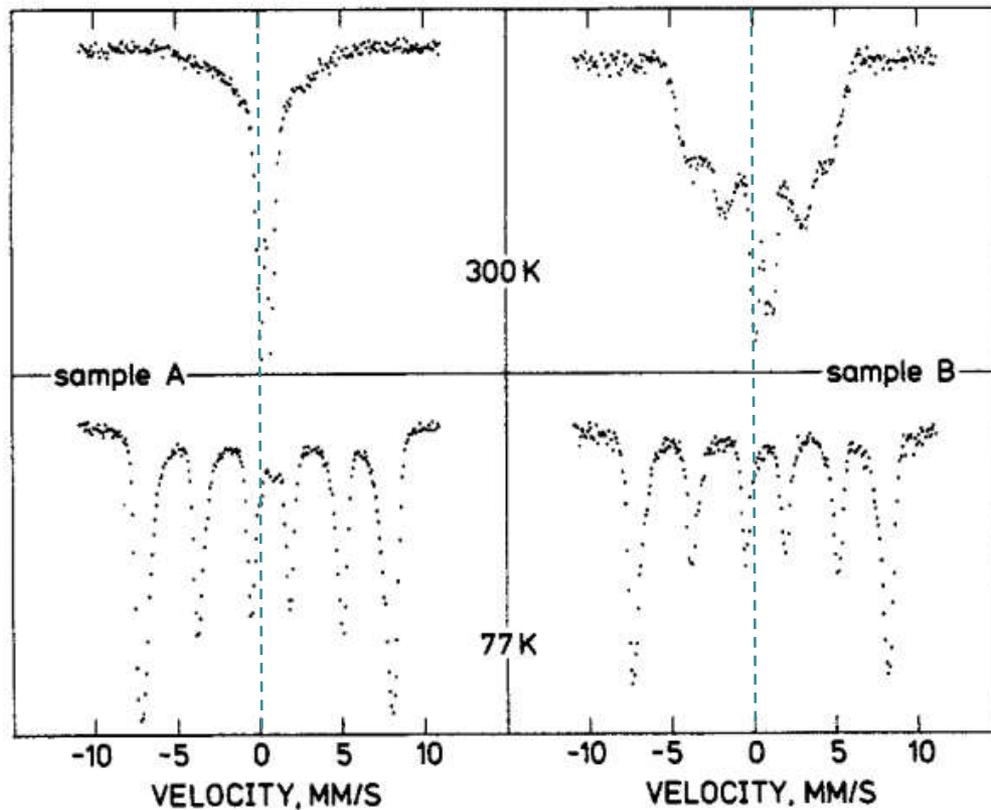
Isomer-shifted Fe hyperfine level scheme and allowed E1 transitions



Mössbauer velocity absorption spectra are shifted against zero and split



Applications in Chemistry/Material Science



Simulation of chemical reaction in Martian (oxygen-free) environment.

Precipitation of Fe in water under influence of UV light:

Intense Hg lamp at distance d from quartz flask with Fe^{2+} solution $d=5\text{cm}$ (A) or $d=10\text{cm}$ (B) \rightarrow Factor 4 in UV intensity

Moessbauer parameters at liquid-nitrogen temperature (77 K) of the photooxidized precipitation samples A and B, compared to natural FeOOH \rightarrow α -FeOOH precipitates (B has larger particles, strong B field at 300K)

| | | δ (mm s ⁻¹) | Δ (mm s ⁻¹) | H (kG) |
|--------------------------------|-----------------|--------------------------------|--------------------------------|----------|
| Photooxidized Fe ³⁺ | sample A | 0.39 | -0.15 | 477 |
| | sample B | 0.38 | -0.11 | 483 |
| Natural FeOOH | α -FeOOH | 0.39 | -0.13 | 487 |
| | γ -FeOOH | 0.49 | 0.54 | 0 |

δ : isomer shift, relative to Fe metal
 Δ : quadrupole splitting
 H : magnetic hyperfine field

Other Mößbauer Cases

^{57}Fe is by far the most common isotope used in Mößbauer experiments.

Isotopes of other elements also frequently studied: ^{129}I , ^{119}Sn , ^{121}Sb

| | | | | | | | | | | | | | | | | | | |
|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|-----------|----|
| H | | | | | | | | | | | | | | | | | | He |
| Li | Be | | | | | | | | | | | B | C | N | O | F | Ne | |
| Na | Mg | | | | | | | | | | | Al | Si | P | S | Cl | Ar | |
| K | Ca | Sc | Ti | V | Cr | Mn | Fe | Co | Ni | Cu | Zn | Ga | Ge | As | Se | Br | Kr | |
| Rb | Sr | Y | Zr | Nb | Mo | Tc | Ru | Rh | Pd | Ag | Cd | In | Sn | Sb | Te | I | Xe | |
| Cs | Ba | La | Hf | Ta | W | Re | Os | Ir | Pt | Au | Hg | Tl | Pb | Bi | Po | At | Rn | |
| Fr | Ra | Ac | | | | | | | | | | | | | | | | |
| | | | Ce | Pr | Nd | Pm | Sm | Eu | Gd | Tb | Dy | Ho | Er | Tm | Yb | Lu | | |
| | | | Th | Pa | U | Np | Pu | Am | Cm | Bk | Cf | Es | Fm | Md | No | Lr | | |

Elements of the periodic table which have known Mössbauer isotopes (shown in red font).



The End